

Figure 1 HYSPLIT forecasts of ash

mass loading for the 2008 eruption of Kasatochi (Aleutian Islands) are produced

using different estimations of the source

meteorological data from a meteorological

model and a source term which is the initial amount of ash. Forecasts are

produced at 5 times which

term. Inputs into HYSPLIT include

Alice Crawford^{1,2}, Barbara Stunder¹, Jaime Kibler³, Michael J. Pavolonis⁴

vent with radius of 100 km

SOURCE TERMS

Compare

1 NOAA/ARL, College Park, MD, 2 UMD/CICS, College Park, MD, 3NESDIS/OSPO/SPSD/SAB-Washington VAAC, College Park, MD

Line - uniform line source above the vent

RT1 – source at location of retrieval at T1

RT2 - source at location of retrieval at T2 RT3 - source at location of retrieval at T3

Meteorological data from a

CYL - uniform cylindrical source centered at

⁴NOAA Center for Satellite Applications and Research, Madison, WI

Figure 7: The top row shows

plots of the satellite retrieved

are the time periods where a

lidar overpass intersected the

Row 2-6 show data from the

CALIOP (Cloud-Aerosol Lidar

with Orthogonal Polarization

data was obtained from the

Atmospheric Science Data

instrument aboard CALIOP. The

NASA Langely Research Center

Row 2 shows curtain plots of

the 532 nm backscatter of the

lidar in shades of blue and gray

and the satellite retrieved top

heights along the lidar track in

Rows 3-6 show the same

curtain plots of the 532 nm

backscatter from the lidar as

row 2 The position of HYSPLIT

computational particles close

to the lidar track are plotted in

Row 2 - Retrieved top heights

Row 3 - Cylindrical source

Row 4 - RT1 GDAS source

Row 5 - RT2 GDAS source

Row 6 - RT3 GDAS source

Dark Blue line- footprint o

165°W

ker, D. M., Z. Liu , A. Omar , J. Tackett, D. Fairlie (2012)

lidar track

https://e

red

red

http://www.arl.noaa.gov/HYSPLIT_info.php

retrievals at each of the 5 time periods (solid lines)

For comparison, CDF's of the modeled clouds which

were initiated with the first satellite retrieval (RT1

ECMWF) are shown (dashed lines). A measure of

how well the modeled and observed mass loading

distributions compare is the absolute value of the

values above the threshold of 0.1 g/m2 are used to

maximum difference between the two CDF's

(Kolmogorov Smirnov Parameter, KSP). CDF's for

calculate the KSP. Because this threshold is not

exact, there is a fairly large error in the KSP. High

values above 0.7 indicate that the distributions ha

little overlap while values below 0.5 indicate fairly

good overlap. Values of KSP are shown in Fig. 5d.

-INTRODUCTION

The HYSPLIT transport and dispersion model is run operationally by the NOAA National Weather Service (NWS) to provide forecast guidance to the NOAA-operated U.S. Volcanic Ash Advisory Centers (VAACs) on the likely evolution of the ash cloud. The model requires meteorological data and an ash source term (position, time, amount and size of ash) as input. Output consists of mass loading or concentration as a function of time and position. Historically, real time quantitative information about an eruption has been sparse. Even estimations of the plume height and duration of the event, two quantities which significantly impact the future spatial extent of the ash cloud, can have significant uncertainties. Consequently the initialization and thus the output of a transport and dispersion model can have large uncertainties. Recently, satellite based volcanic ash retrieval algorithms have been developed which can identify the spectral signature of ash and then calculate the mass loading, cloud height and effective radius as a function of position for the ash cloud. This type of information can dramatically reduce uncertainty in model output by reducing uncertainty in the source term. It also has the potential to direct model development by providing more opportunities for model evaluation. Here we look at the 2008 eruption of Kasatochi in the Aleutian Islands. We use satellite retrievals provided by NOAA/CIMSS for this eruption which utilized data from the MODIS instrument. We use five retrievals which are spaced 11 to 12 hours apart. The first retrieval occurred near the end of the main eruption. We compare HYSPLIT output produced using different source terms. We use the satellite retrievals both to help construct some of the source terms and for verification. For further verification, data from the CALIOP lidar instrument aboard CALIPSO is used. This case illustrates the potential for improving transport and dispersion model output by using satellite information to construct source terms and for model evaluation

HORIZONTAL EXTENT OF ASH



Figure 3: The cumulative distribution function (CDF) of the mass loading produced by the cylindrical source at all five time periods. In this example the mass loading of the modeled cloud spans 5-6 orders of magnitude. For HYSPLIT runs initiated at the vent, the model is configured to emit one unit mass /h and the output is given in unit mass/m². A threshold must be chosen in order to compare the spatial overlap of the modeled and observed clouds. Four thresholds are chosen as shown. The unit mass can be converted to grams by estimating that the threshold is equal to 0.1 g/m². The light blue shaded area indicates a range of thresholds which might be chosen by applying an empirical equation with a plume height of 14-18 km and a mass fraction of fine ash of 10%. This type of graph illustrates the effect of changing the threshold. The different thresholds are equivalent to different mass eruption rates (MER) or to different ash reduction levels available operationally or on the web in the VAFTAD format.



Figure 5: The CSI, POD, FAR and KSP are plotted for different source terms and meteorological data sets for the 5 time periods. The RT1 source consistently produced the highest CSI by having a fairly large POD and one of the lowest FARs. RT2 and RT3 had similar CSI, but there appears to be no advantage to using the latest satellite retrieval for initialization in this example. The Line source (same threshold as CYL A) initially produced a plume that was too narrow (low POD and FAR), however the plume spread with time. Although CYL B does well for the first time period, the plume dissipates much too quickly (low FAR and POD). CYL A, C and D had similar CSI's. CYL C and D had both the highest POD's and highest FAR's. CYL B , C and D have a very high KSP, indicating that the mass distribution has little overlap with the observed mass distribution. For the sources initiated at the vent, picking a threshold value of 1x10⁻¹² mass units / m² (CYL A), reproduces the observed mass loading the best.





Using a source term located at the observed position of the ash cloud produces better or comparable results to using a source term located at the vent. The mass loading retrieval can reduce uncertainty in the forecast modeled ash cloud by providing better information on which portions of the modeled cloud would be too diffuse to be detected. A cylindrical source may represent the early plume better than a line source for large eruptions.

icknesses produced comparable sults to the ones shown.



T1 - 2008 08/08

13:40 UTC

T2 - 2008 08/09

00:50 UTC

Figure 2: Passive IR satellite retrievals² of

column mass loading of ash at five time periods.

-143.71 10 11:5 45.91 -143.71 10 11:5 44.1 55.90 -166.20 10 00:01 -163.41 09 23:59 -159.20 -161.13 09 23:57 -141.30 10 11:52 -134.33 10 11:49 -138.31 10 11:51

METHOD

12:50 UTC

HYSPLIT output is able to capture the vertical structure of the ash cloud well. The output using the RT2 and RT3 sources do not agree as well with the lidar data at T4 as the output using the RT1 source. Lidar data, retrieved top heights, and model output all suggest that the ash cloud evolves into a complex three dimensional structure with patches and layers. The RT1 source may have an advantage over the RT2 and RT3 sources in that the early cloud has a simpler structure and is better represented by a single layer.

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135°V

swept up by a low pressure system. The presence of this

strong synoptic scale feature

modeled cloud at T2-T5 less

sensitive to initial conditions

CSI is lower at T1 than at

not be modeled as well.

and is probably why we see the

subsequent times (T2-T4). At T1 the cloud shape is influenced

by the eruption dynamics and smaller scale winds which may

almost certainly makes the